# Three-dimensional nature of magnetic reconnection X-line in asymmetric current sheets

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# Three-dimensional nature of magnetic reconnection in plasmas

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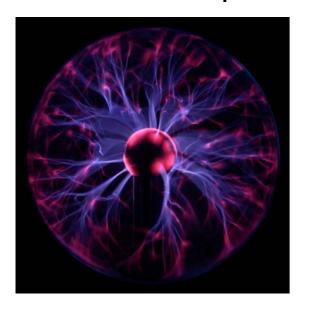


Key Challenge & Why it matters?

**Plasmas** 

4th state of matter
> 99% of visible universe
Fusion device

Plasma Lamp

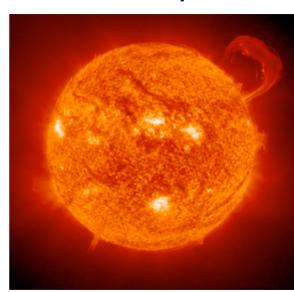


Lighting

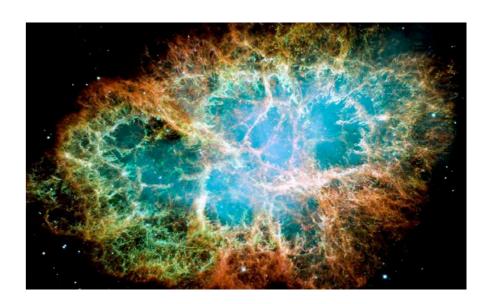
Aurora Borealis



Solar Eruption



Nebula

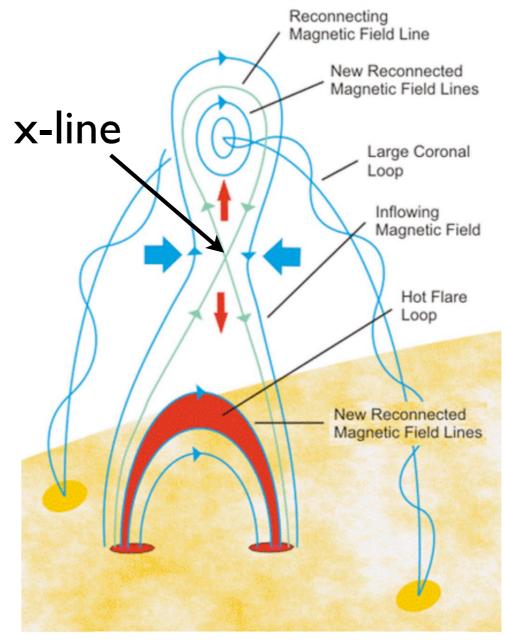


- Interaction between lots<sup>n</sup> of charge particles + electromagnetic fields
   -- complicated & nonlinear!!
- Long range electromagnetic interaction!!
  - -- the evolution CANNOT be described by thermodynamics.

# Introduction to magnetic reconnection -- Solar Eruption



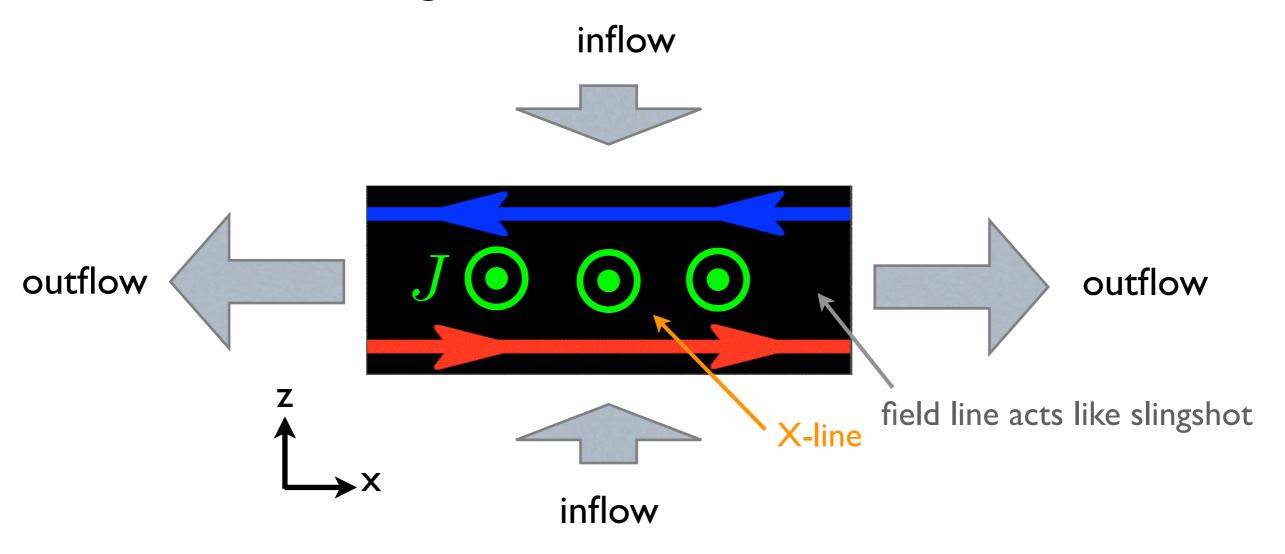
(Observation of SDO mission) B~200 Gauss T~3,000,000 K



(Courtesy of NASA)

- Energy up to 10<sup>32</sup> ergs is released in ~ 20 mins
   40 billion atomic bombs!
- Matter up to 10<sup>10</sup> tons is erupted.

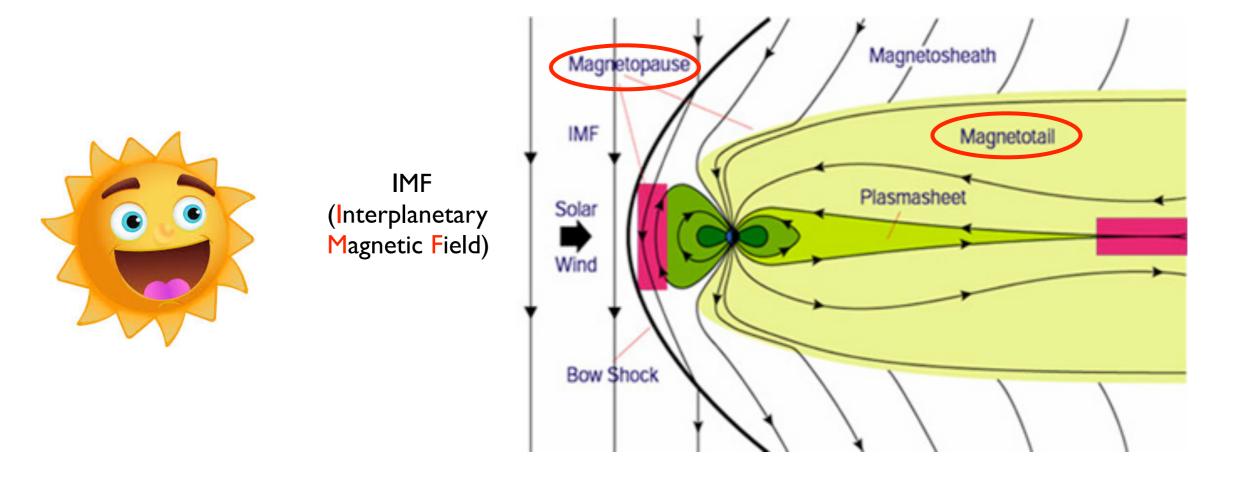
### Magnetic Reconnection?



- 1. Inflow brings in magnetic flux
- 2. Field lines break & reconnect
- 3. Reconnected field line shoots out plasma
- 4. Pressure drop sucks in plasma inflow
- I. Inflow brings in magnetic flux
- 2. ....
- 3. ...

A self-driven process!!!

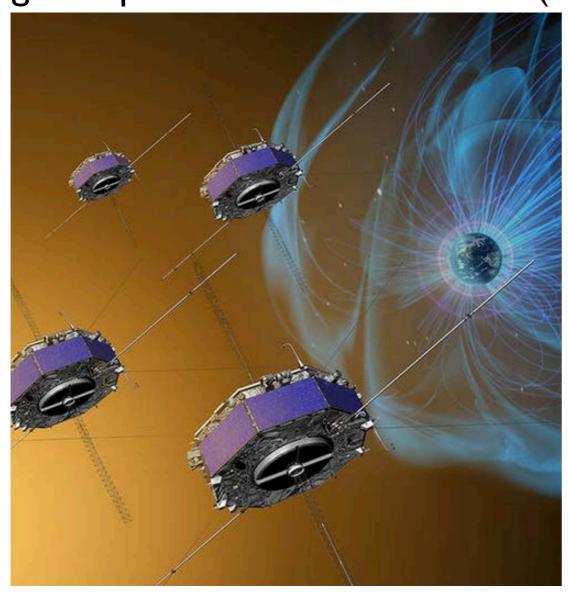
### Earth's magnetosphere



- Reconnection occurs at both the magnetopause & magnetotail.
- Reconnection at the magnetotail drives magnetospheric substorm
   & causes aurora.
- Space Weather: a strong geomagnetic storm could do damage to satellites, astronauts, GPS system, power grids on Earth,....etc (e.g., 1859, 1989) Quebec blackout)

# A billion \$ NASA mission designed to study magnetic reconnection

Magnetospheric Multiscale Mission (MMS)





http://mms.gsfc.nasa.gov

tight tetrahedron formation: separation down to 7 km! 100x faster for electrons measurement (30 ms) 30x faster for ions measurement (150 ms)

• MMS leads us into a stage where the 3D electron-scale structure of magnetic reconnection, in nature, can be measured in an unprecedented manner.

# Why Blue Waters? & Our accomplishments to date

### Particle-in-cell Simulations

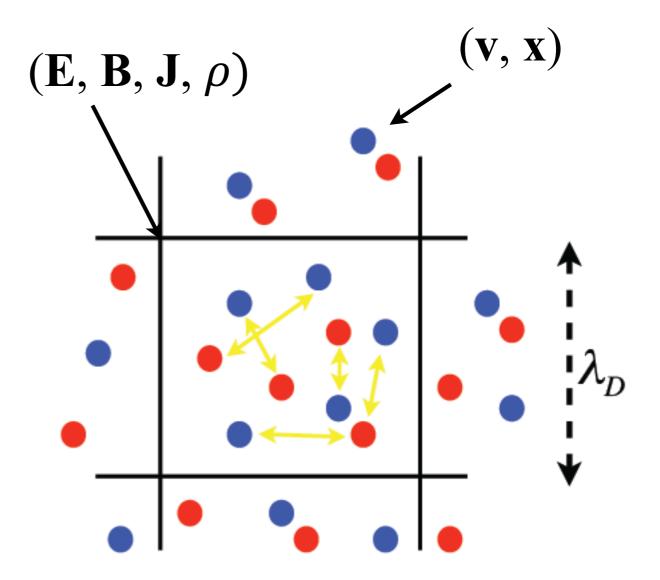
#### **Lorentz Force**

$$\mathbf{F} = q(\mathbf{E} + \mathbf{v} \times \mathbf{B})$$

### Maxwell Equation

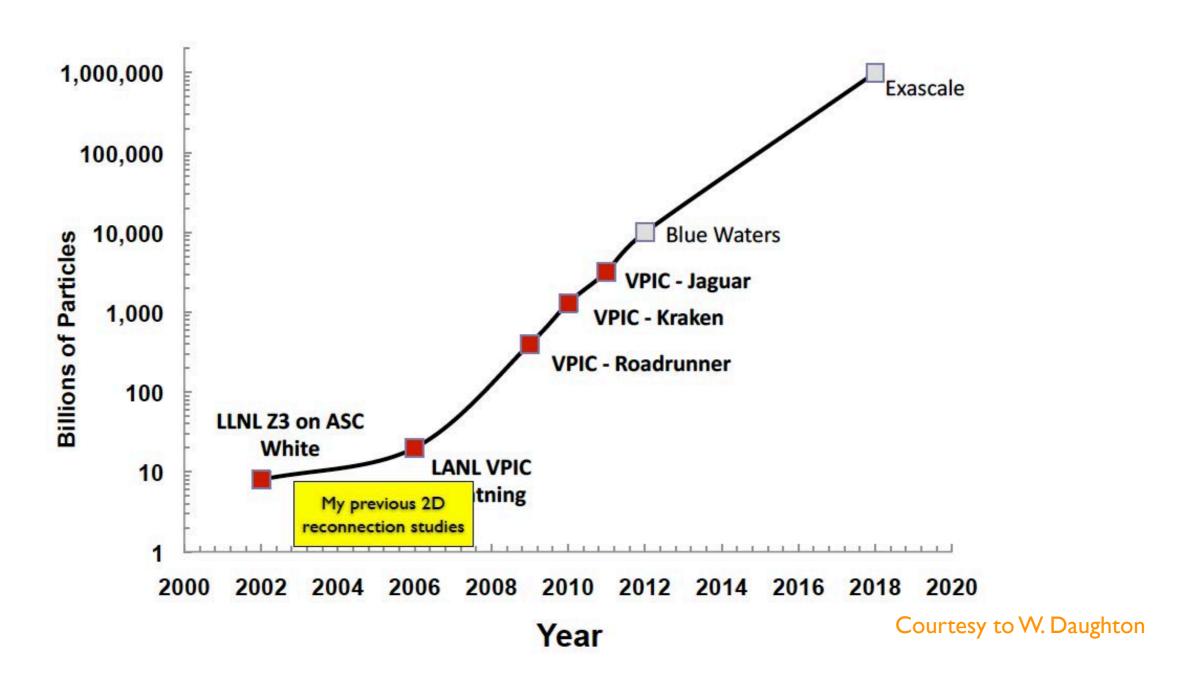
$$\nabla \cdot \mathbf{B} = 0 \qquad \nabla \times \mathbf{B} = \frac{4\pi}{c} \mathbf{J} + \frac{1}{c} \frac{\partial \mathbf{E}}{\partial t}$$

$$\nabla \cdot \mathbf{E} = 4\pi \rho \quad \nabla \times \mathbf{E} = -\frac{1}{c} \frac{\partial \mathbf{B}}{\partial t}$$



Pro: A first-principle description with rich kinetic physics being captured <a href="Con">Con</a>: It demands considerable computational resource ,especially for 3D systems (Challenge!)

## Progress of Particle-in-Cell simulations

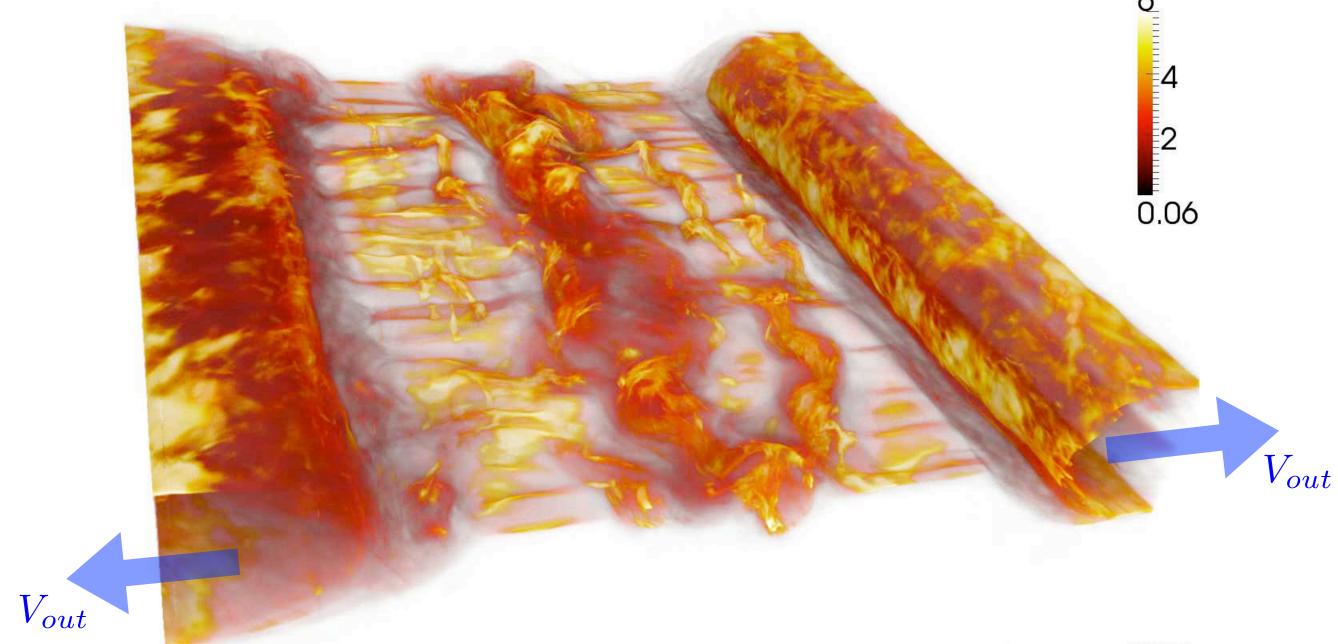


- Resource required for a typical run in our BW project:
  - ~2 trillion particles; ~ 6 billion cells;
  - ~12 million core-hours; ~ 260 K cores;
  - ~200 TB output (including restart files)

# Q: What is the 3D nature of magnetic reconnection?? -- how it differs from the conventional 2D picture

## An example 3D simulation

 $\begin{array}{c} \text{(current)} \\ \text{density)} \\ J/J_0 \end{array}$ 



Distinct 3D features, including

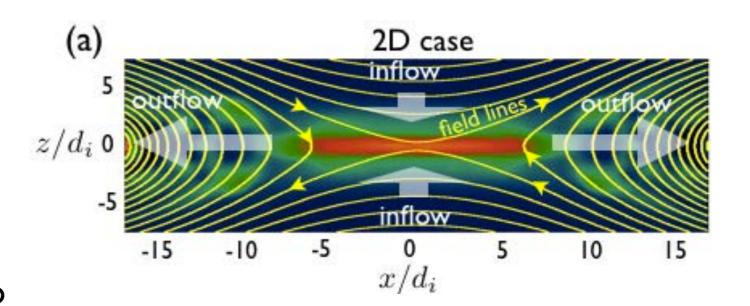
- flux ropes.
- kink instability.
- turbulence.

$$t\omega_{pe} = 382$$

(Previous BW results. Guo, Li, Daughton, Liu, PRL 2014 Guo, Liu, Daughton, Li, APJ 2015)

### 3D nature of reconnection X-line

Q:What determines the orientation of x-line??



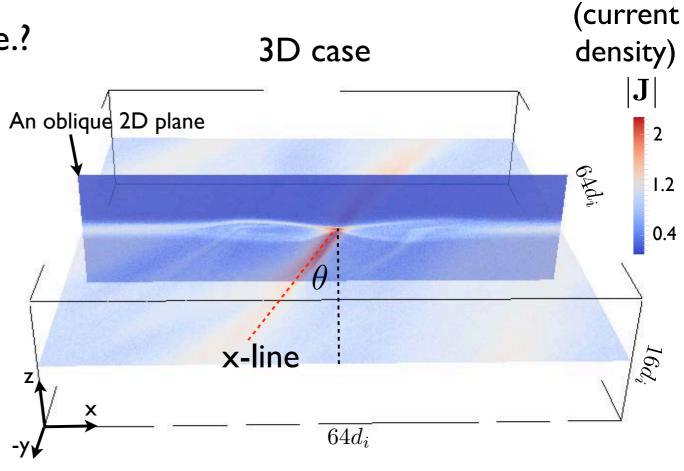
Hypothesis: the system may tend to maximize

-- the magnetic energy conversion rate.?

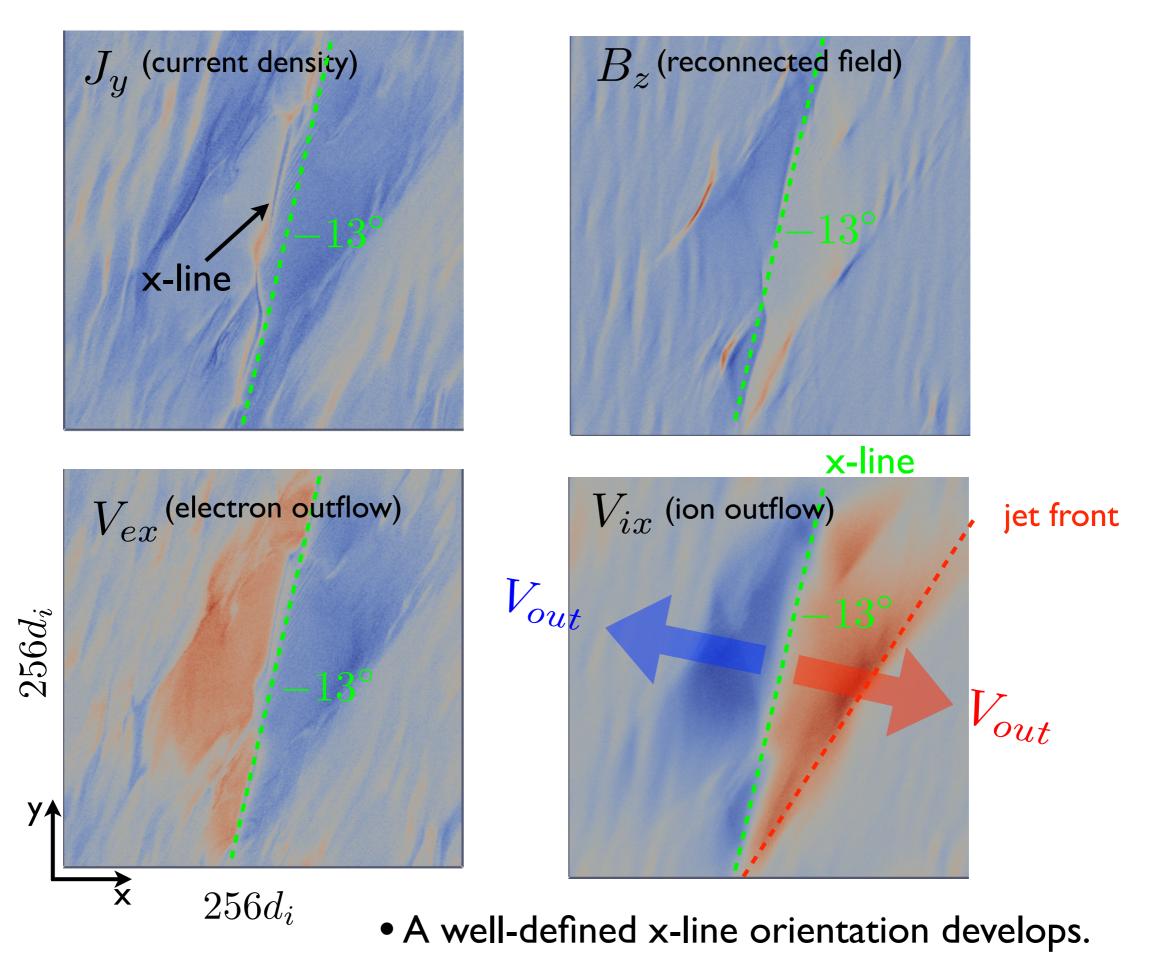
-- the outflow speed.?

-- or ??

 Simulation on BW provides a first-principle test!

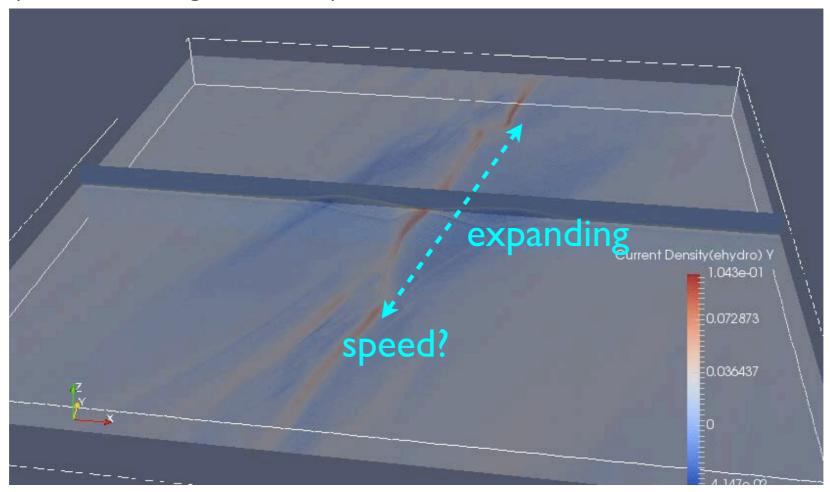


Result I: X-line Orientation



# Result II: Extent and expansion of reconnection x-line in the magnetopause geometry

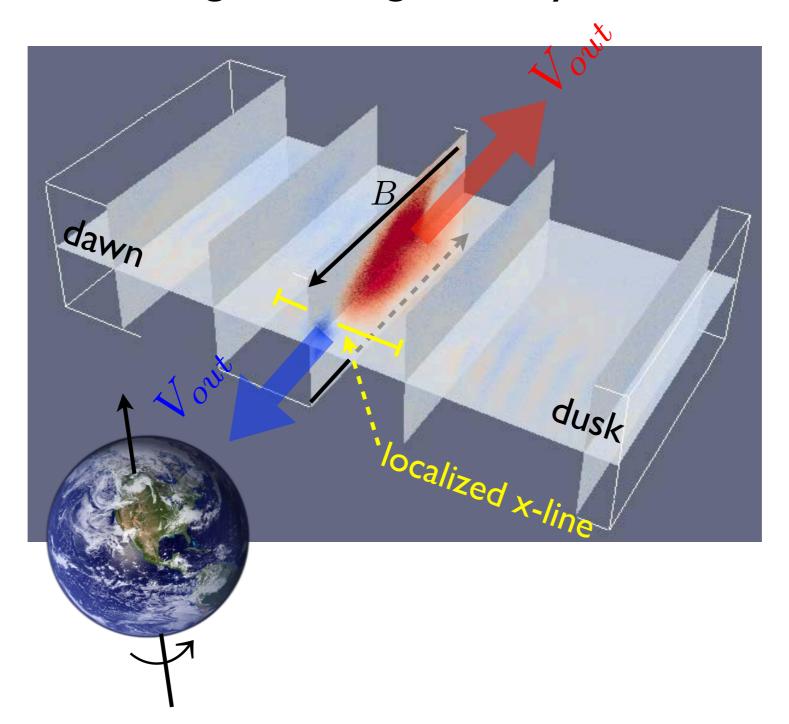
(3D view using ParaView)

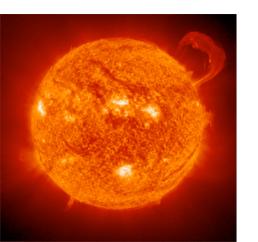


Q: How does the x-line expand? what is the expanding speed?

--- Knowing the length of x-line is crucial for estimating the entry rate of solar wind plasmas into Earth's magnetosphere.

## Result III: Extent and expansion of reconnection x-line in the magnetotail geometry



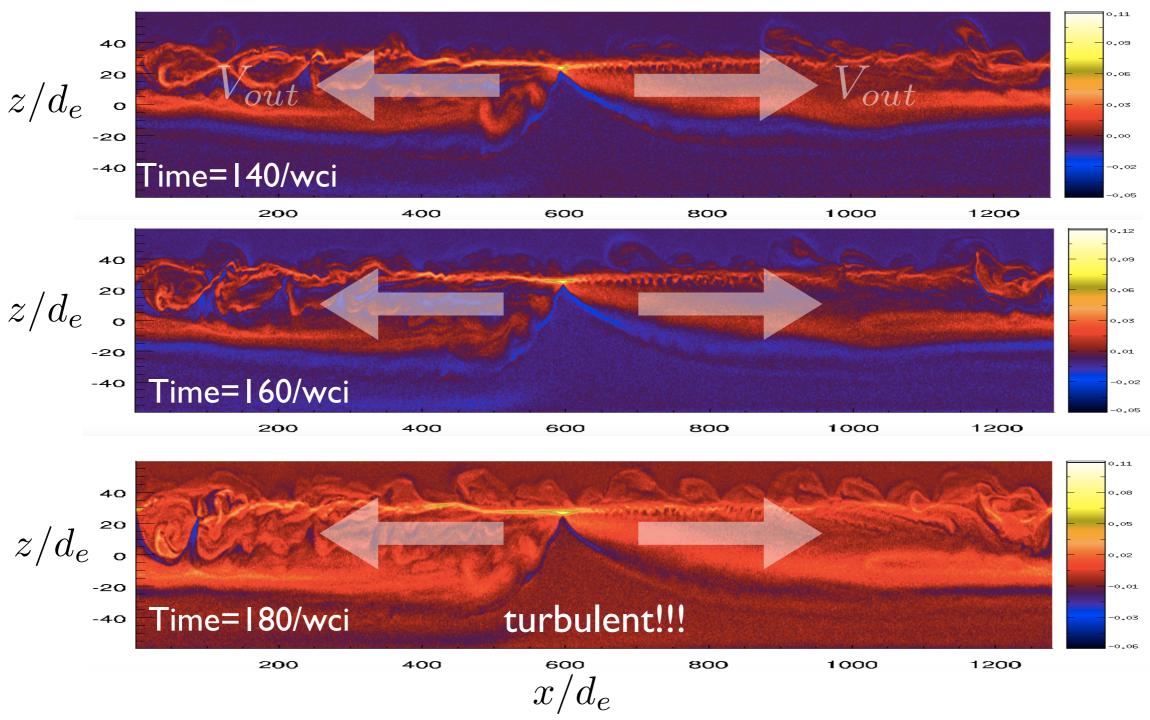


Q: Is there a limited x-line extent in the tail geometry?

-- Relevant to the formation of Bursty Bulk Flows (BBFs) often observed at Earth's magnetotail.

### Result IV: Self-generated turbulence in Reconnection

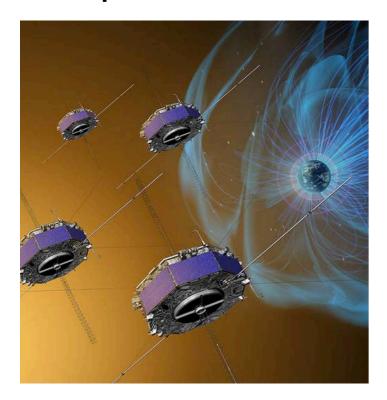
## Current density



- The reconnection current sheet & x-line can become turbulent in 3D!
  - -- the cause? the effect on the dissipation & energy conversion rate?

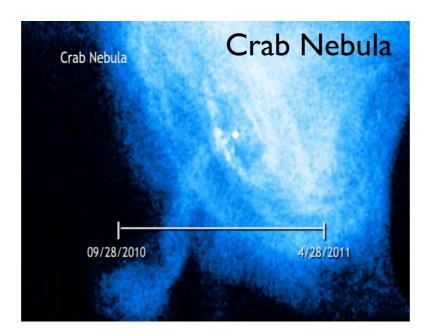
### Broader Impact

### Space Science



NASA's MMS mission

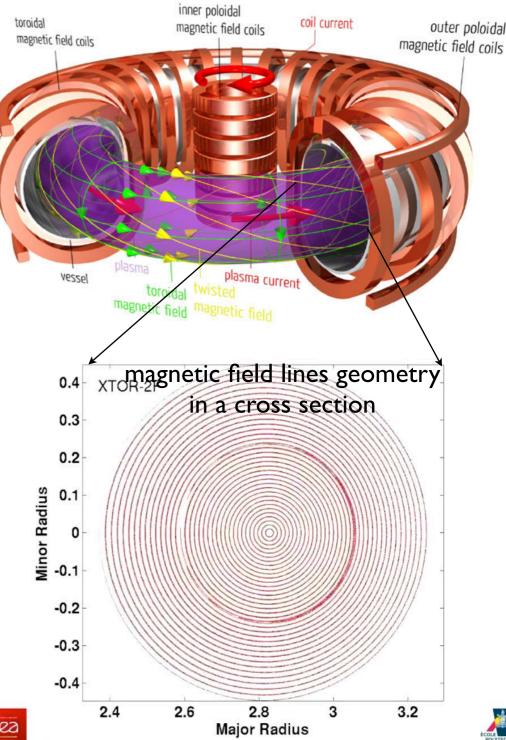
### **Astrophysics**



Fermi Gamma Ray Space Telescope

#### Fusion device

#### e.g., ITER Tokamak @ France







### Blue Waters Team Contributions

### Special thanks to

Ryan Mokos for the help & advices on the data archiving.

JaeHyuk Kwack for the script that simultaneously launches tar comment in multiple cores!

Craig Steffen & storage group for fixing the error in my file system at Nearline.

Roberto Sisneros & visualization group for helping setup the 3D visualization using ParaView.

## Summary

Magnetic Reconnection is the key process that releases the magnetic energy stored in space, astrophysical and laboratory plasma systems.

- -- The 3D nature of reconnection x-line remains unclear!
- -- Blue Waters provides the opportunity to explore this challenging problem.

The on-going science topics run on Blue Waters include:

- -- The orientation of 3D x-lines.
- -- The extent and expansion of 3D x-lines with local geometries similar to Earth's magnetopause & magnetotail.
- -- The turbulent nature of 3D x-lines.